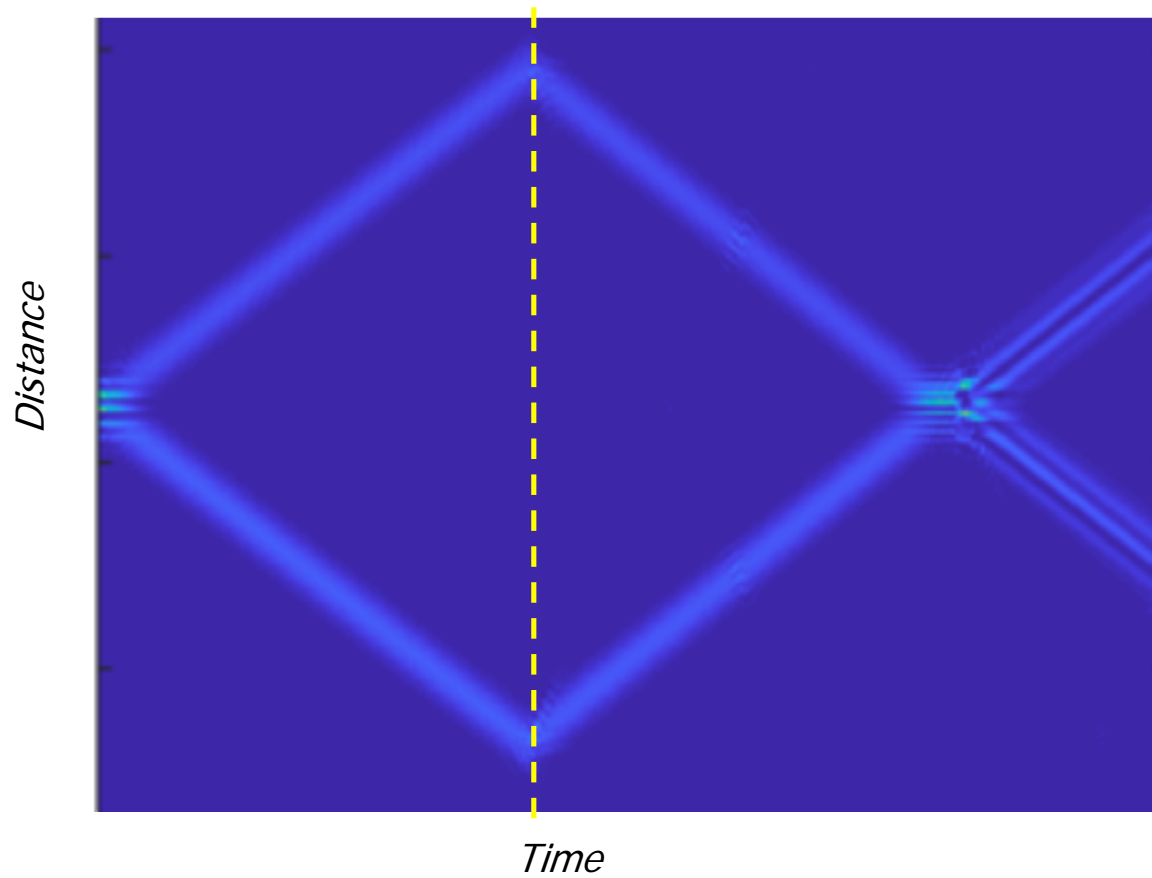


# Tests of Gravity and Quantum Mechanics with Atom Interferometry

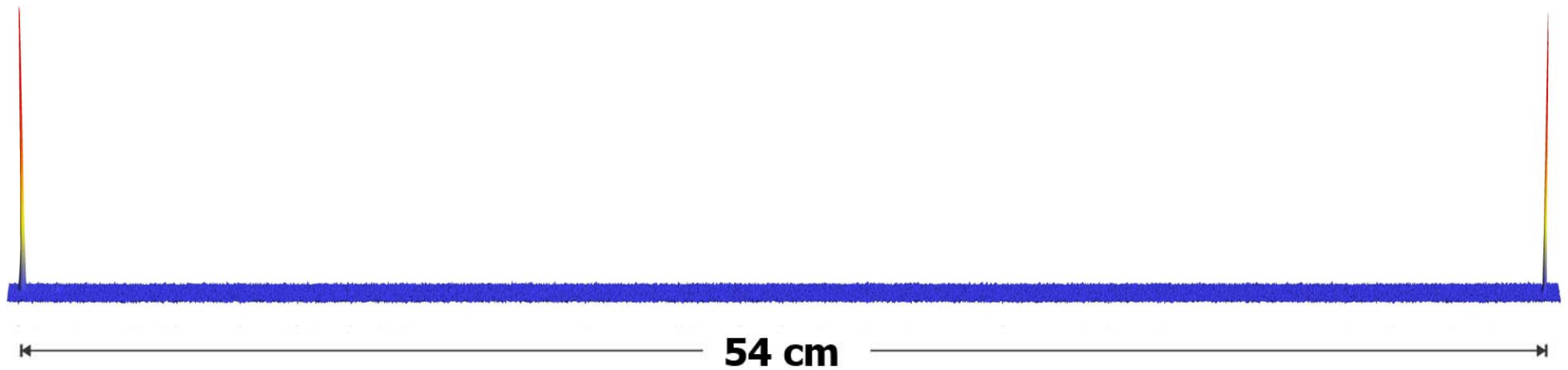
Mark Kasevich  
Stanford University  
kasevich@stanford.edu



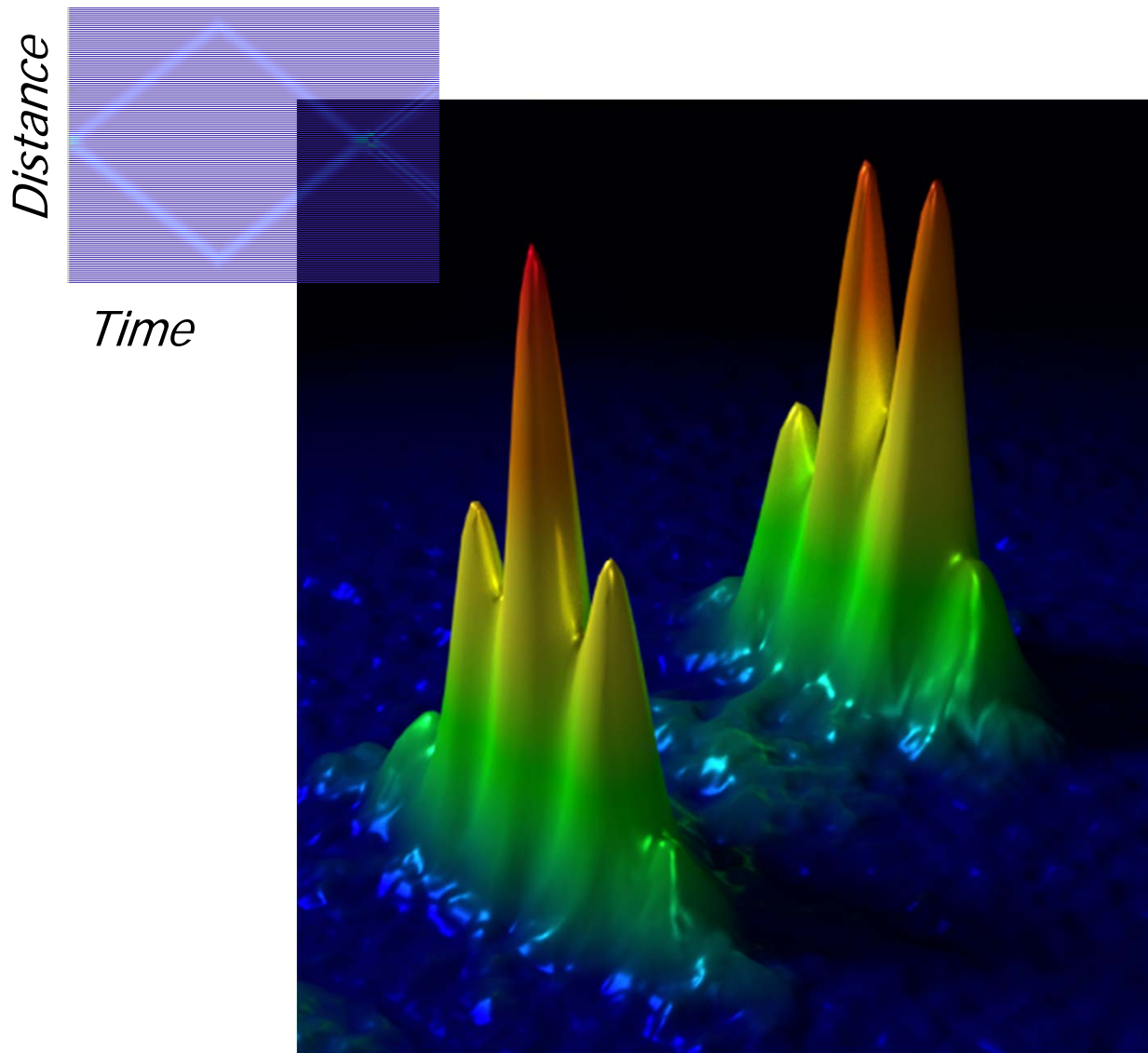
# Atom interferometry



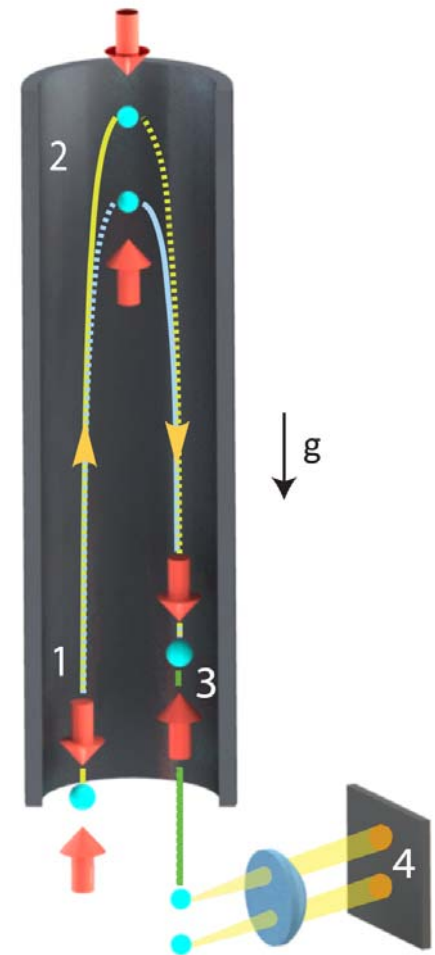
# Atomic wavepacket superposition



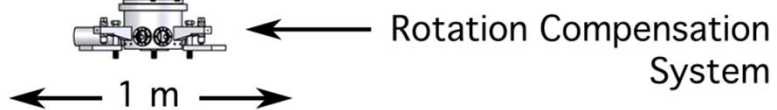
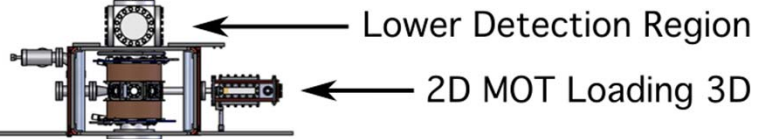
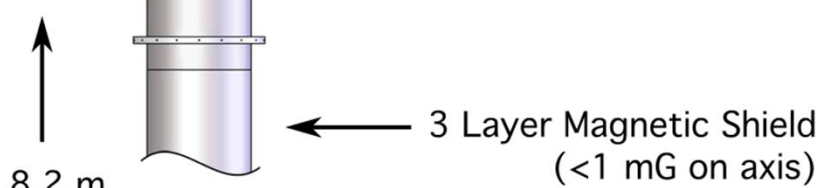
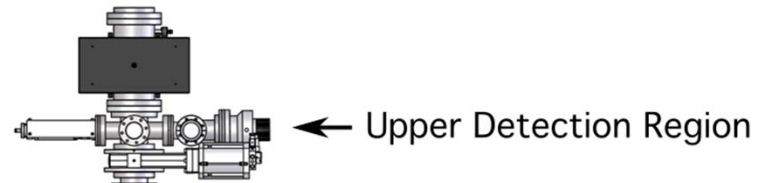
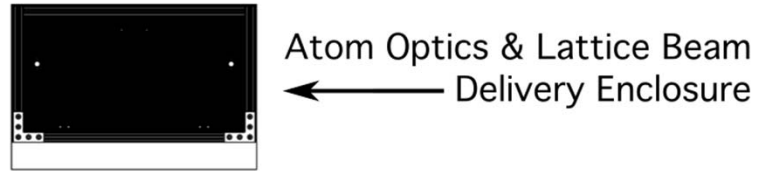
# Atomic wavepacket interference



*Superposed atomic wavepackets interfere*



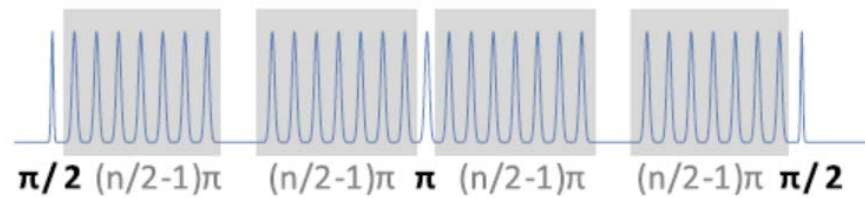
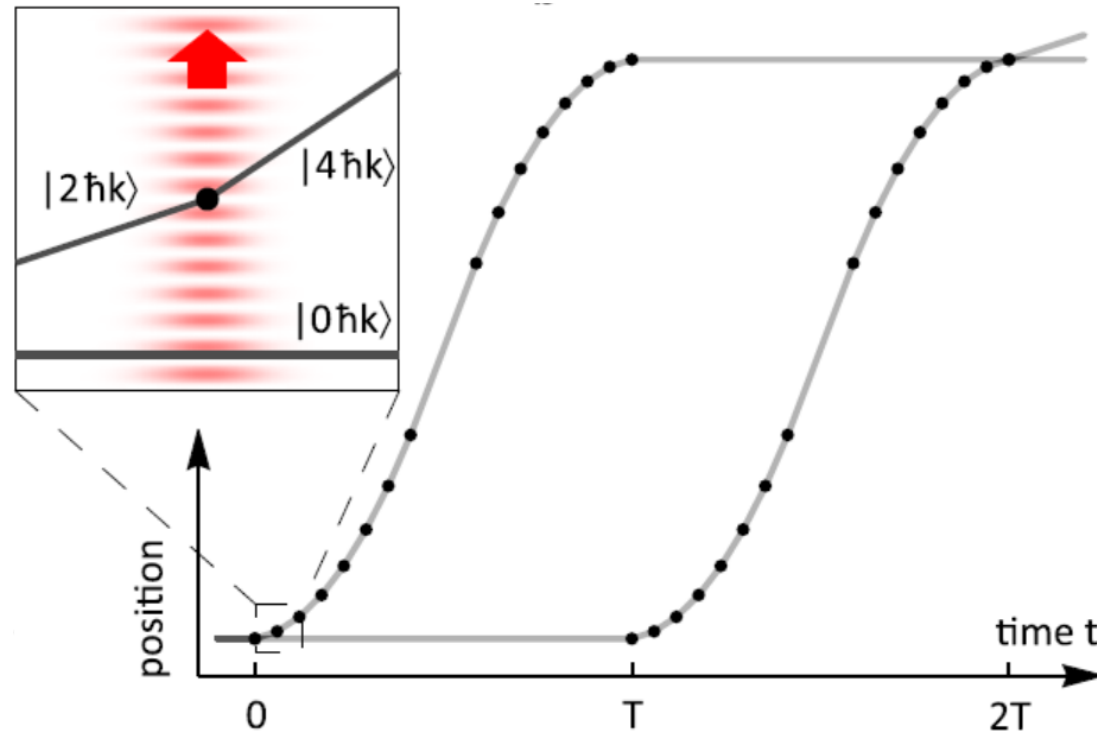
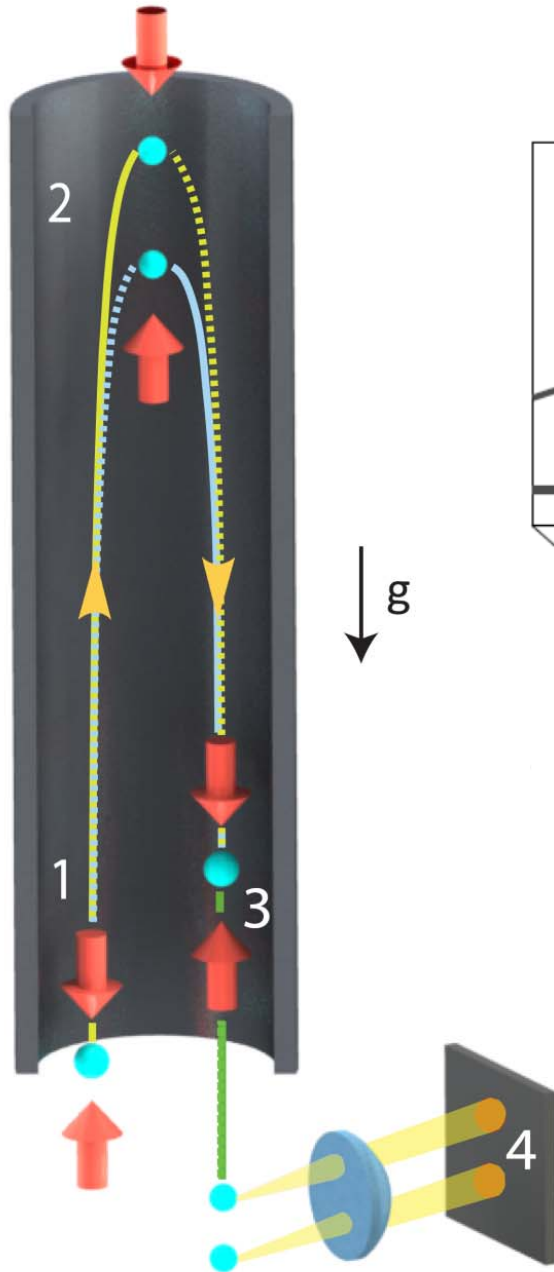
# Atomic fountain



$\sim 100$  pK  
 $1e5$  atoms/shot



# Sequential Bragg atom optics





Pulse sequence duration: 2.08 s

# Bounds for KTM model

## Gravity is not a pairwise local classical channel

$$\tilde{\Gamma}_{\text{KTM}}^C = C \frac{GMm}{\hbar R^3} \Delta x^2$$

Natacha Altamirano<sup>1,2</sup>, Paulina Corona-Ugalde<sup>2,3</sup>,  
Robert B Mann<sup>1,2,3</sup>  and Magdalena Zych<sup>4,5</sup> 

Experiment	m (Kg)	M (Kg)	d (m)	$\Delta x$ (m)	$1/\Gamma_{\text{DP}}$ (s)	$1/\Gamma_{\text{KTM}}^{\text{min}}$ (s)
10 m atomic fountain with <sup>87</sup> Rb [38]	$1.4 \times 10^{-25}$	$M_{\oplus}$	$R_{\oplus}$	0.54	$3 \times 10^{10}$	$2 \times 10^{-3}$
Two atomic fountains with <sup>87</sup> Rb [33] (Operating as gravity-gradiometer)	$1.4 \times 10^{-25}$	$M_{\oplus}$	$R_{\oplus}$	$1.86 \times 10^{-3}$	$3 \times 10^{10}$	$2 \times 10^1$
		$4 \times 129$	0.11, 0.18, 0.28, 0.31			
Large-molecule interferometry [43]	$1.6 \times 10^{-23}$	$M_{\oplus}$	$R_{\oplus}$	$2.7 \times 10^{-7}$	$3 \times 10^6$	$6 \times 10^7$
PcH <sub>2</sub> diffraction on alga skeleton [44]	$8.2 \times 10^{-25}$	$M_{\oplus}$	$R_{\oplus}$	$2 \times 10^{-7}$	$1 \times 10^9$	$2 \times 10^9$



# Phase shifts (non-relativistic)

	Term	Phase Shift	
8e9 rad →	1	$k_{\text{eff}} g T^2$	Gravity
	2	$2\mathbf{k}_{\text{eff}} \cdot (\boldsymbol{\Omega} \times \mathbf{v}) T^2$	Coriolis
	3	$k_{\text{eff}} v_z \delta T$	Timing asymmetry
	4	$\frac{\hbar k_{\text{eff}}^2}{2m} T_{zz} T^3$	Curvature, quantum (tidal)
635 rad →	5	$k_{\text{eff}} T_{zi} (x_i + v_i T) T^2$	Gravity gradient
	6	$\frac{1}{2} k_{\text{eff}} \alpha (v_x^2 + v_y^2) T^2$	Wavefront

$T_{ij}$ , gravity gradient

$v_i$ , velocity;  $x_i$ , initial position

$g$ , acceleration due to gravity

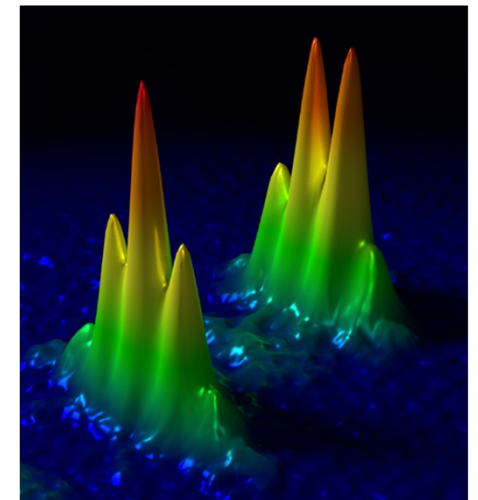
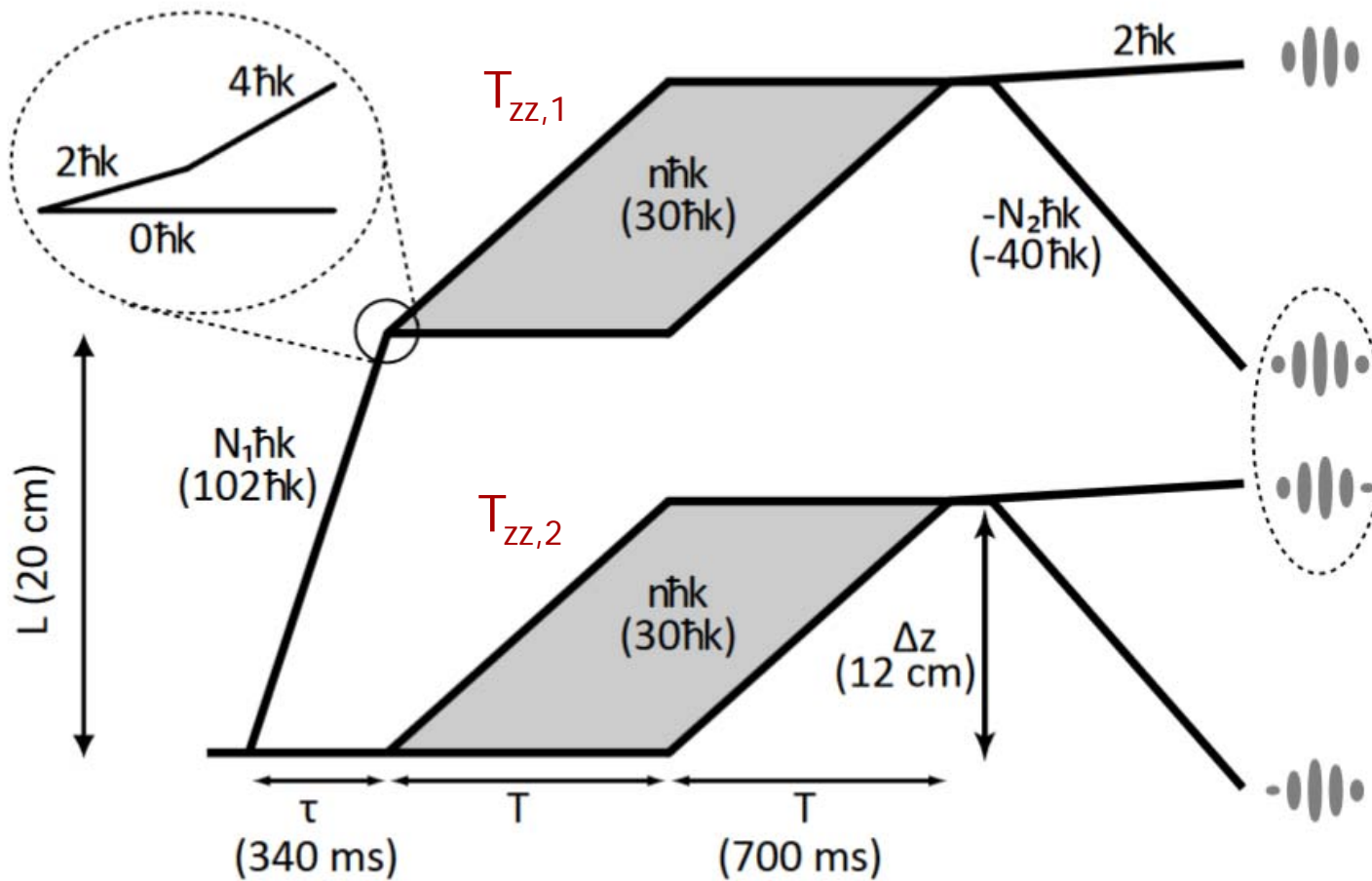
$T$ , interrogation time

$\mathbf{k}_{\text{eff}}$ , effective propagation vector





# Curvature, quantum (tidal) phase shift observation

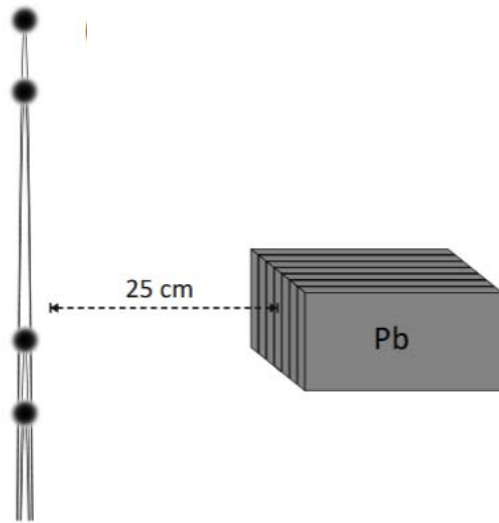


Resolution:  $2e-9/s^2$   
per shot  
( $7e-4$  Earth  
gradient/shot)

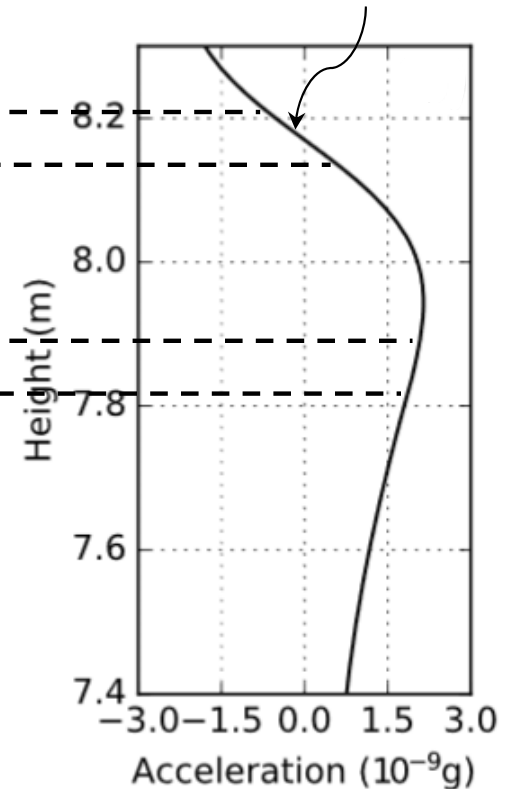
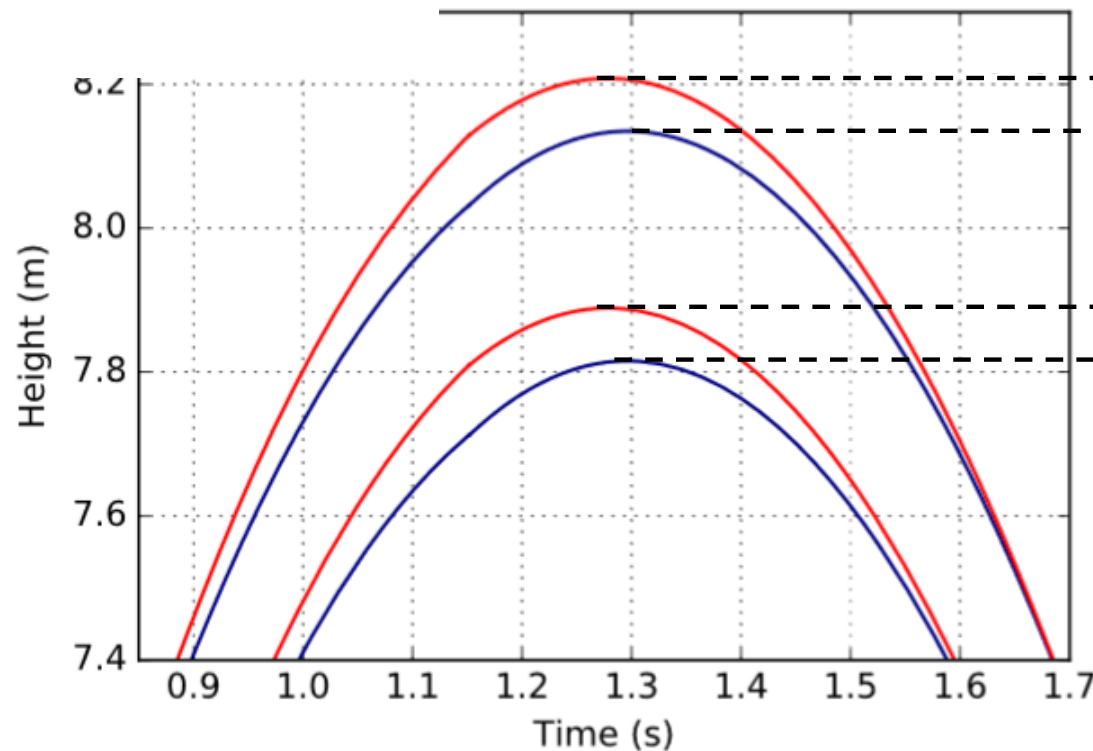
Use gravity gradiometer configuration to isolate quantum curvature phase shifts

# Tidal forces on a wavefunction

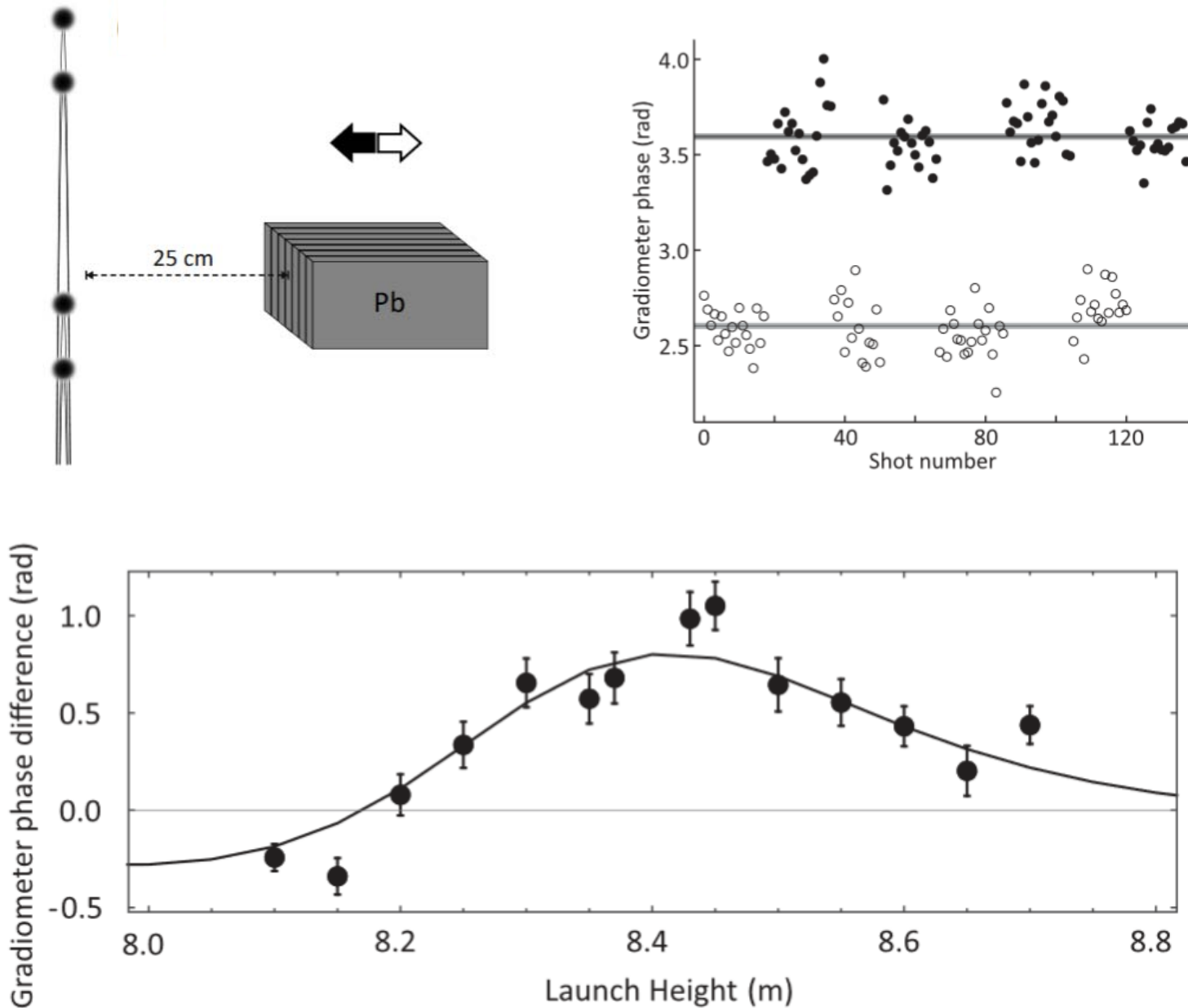
Exploit large wavepacket separations to directly observe the influence of gravitational tidal forces on a wavefunction.



Each interferometer arm experiences a (resolvable) different force

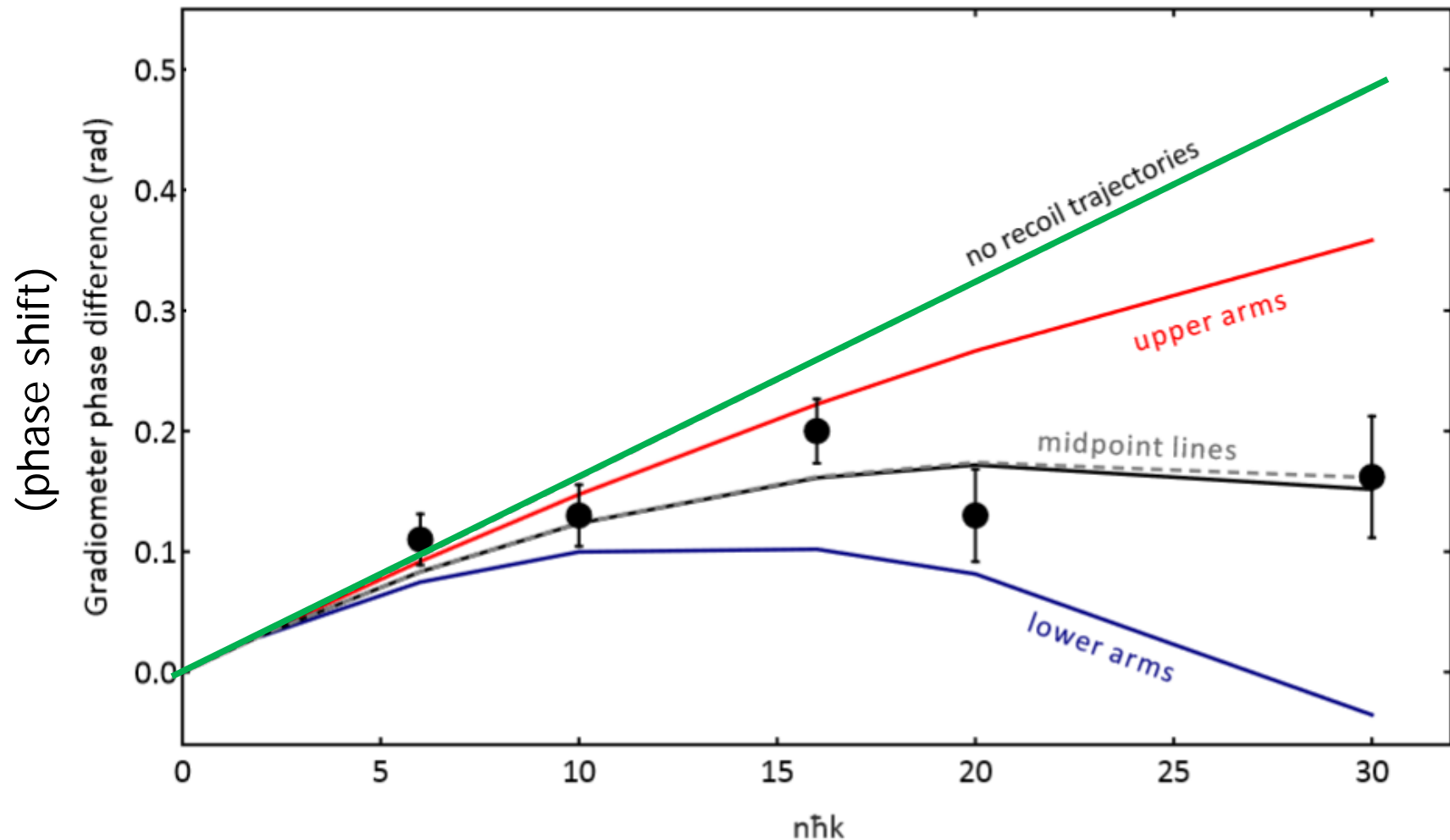


# Observation of tidal forces on a wavefunction



# Gravity and quantum mechanics

First observation of the influence of gravitational curvature on any quantum system (see Audretsch, PRA 1994; Anandan, PRD, 1984).



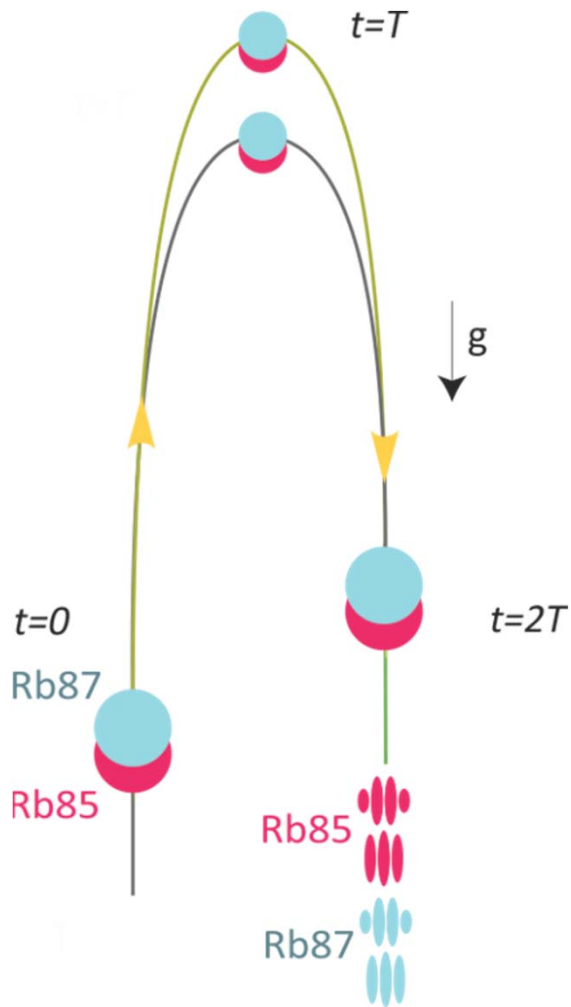
Asenbaum, PRL (2017)

STANFORD UNIVERSITY

(wavepacket separation)



# Equivalence Principle Test

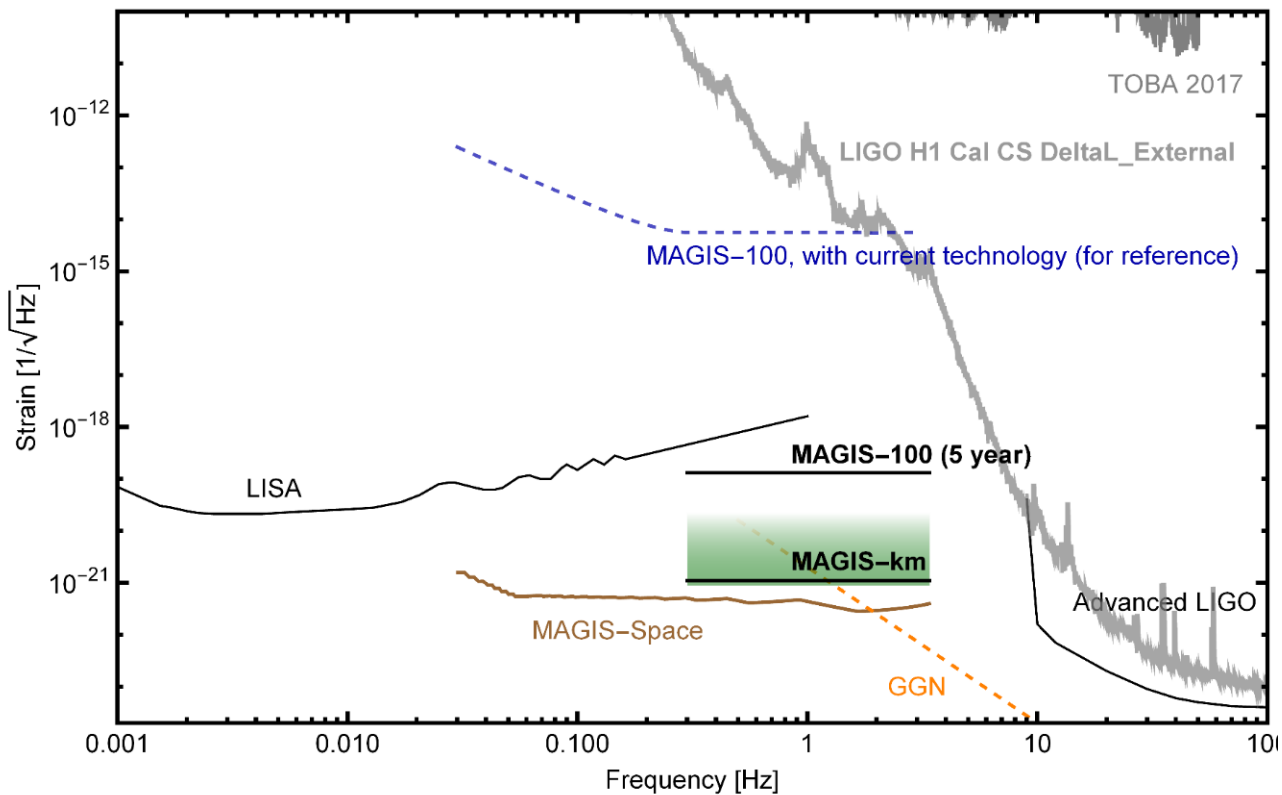
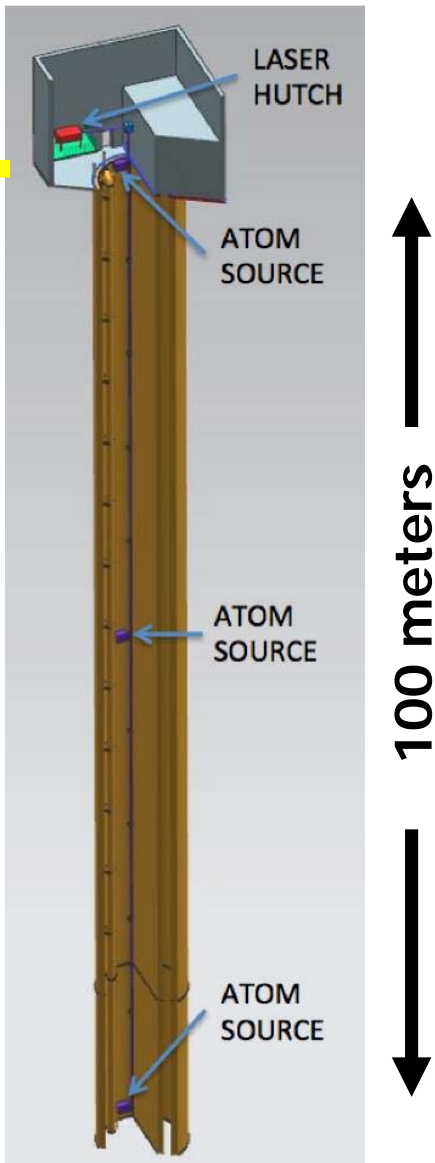
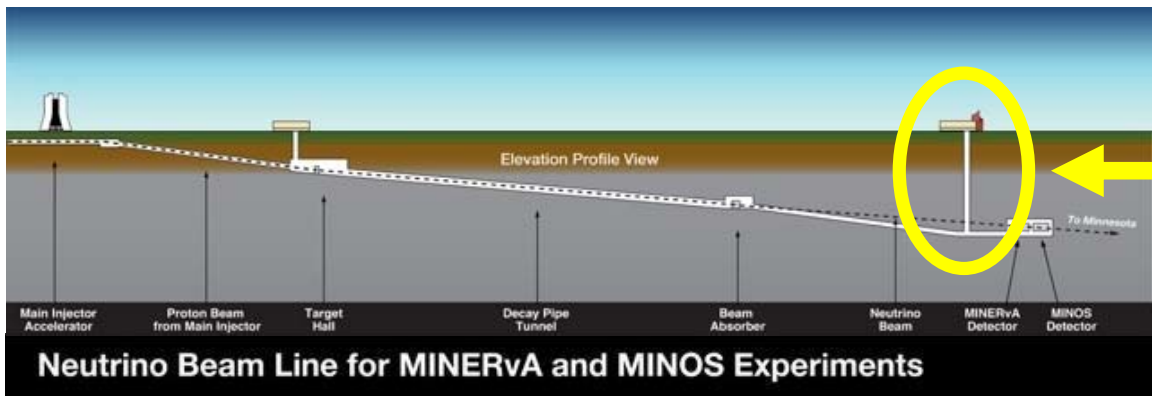


85 Rb

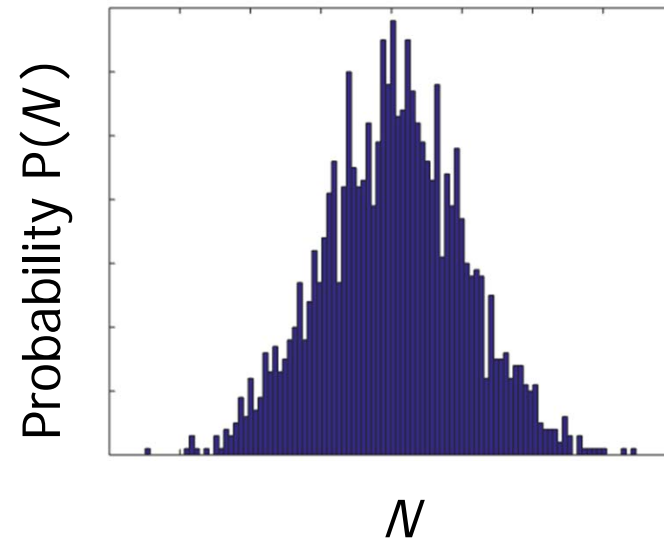
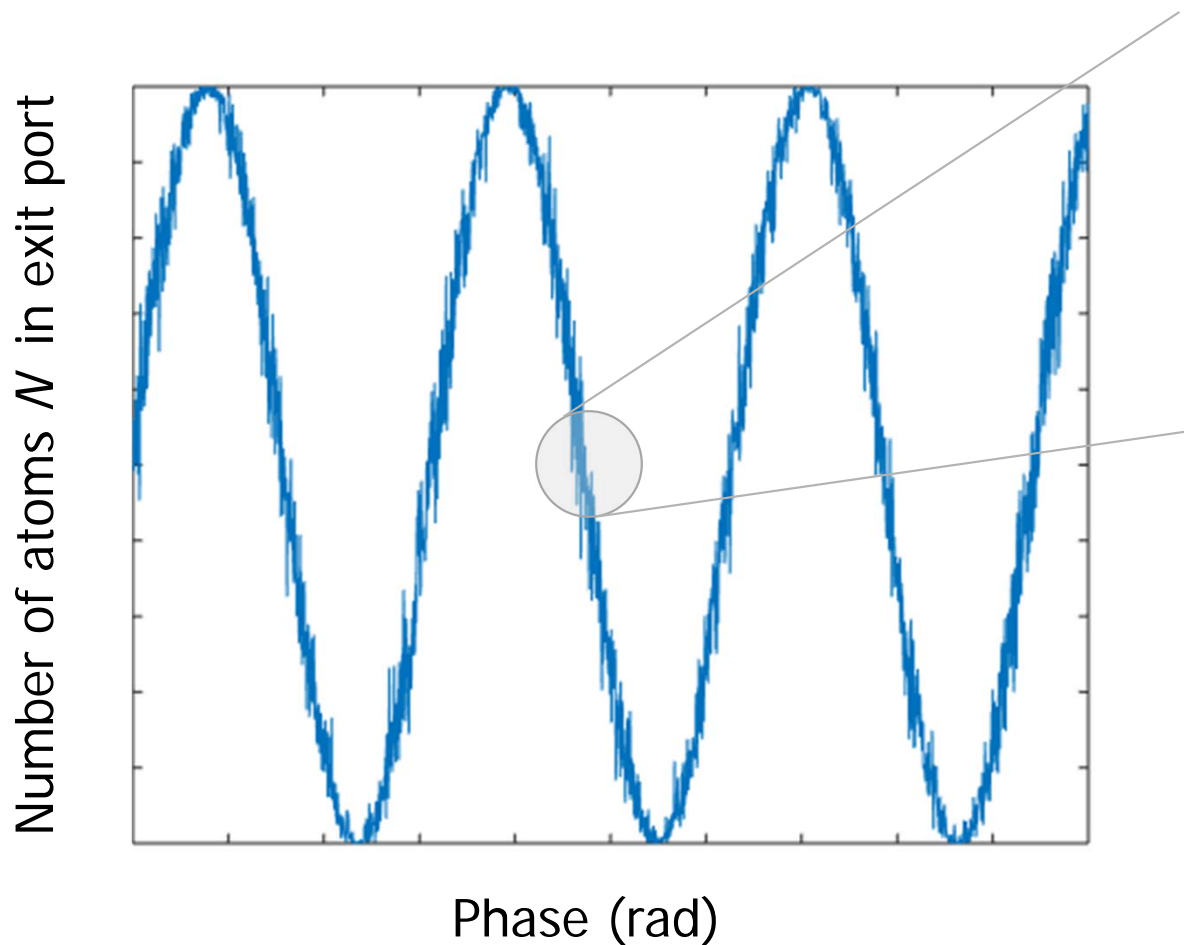
87 Rb

Ground-based EP (target precision  $\sim 1e-14 \delta g/g$ )  
5.4 cm wavepacket separation

# MAGIS-100: GW detector prototype at Fermilab



# Noise in interferometric sensors



For uncorrelated particles:

“Shot noise” – physical origin is randomness of wavefunction collapse.

# Quantum correlated (entangled) atomic ensembles

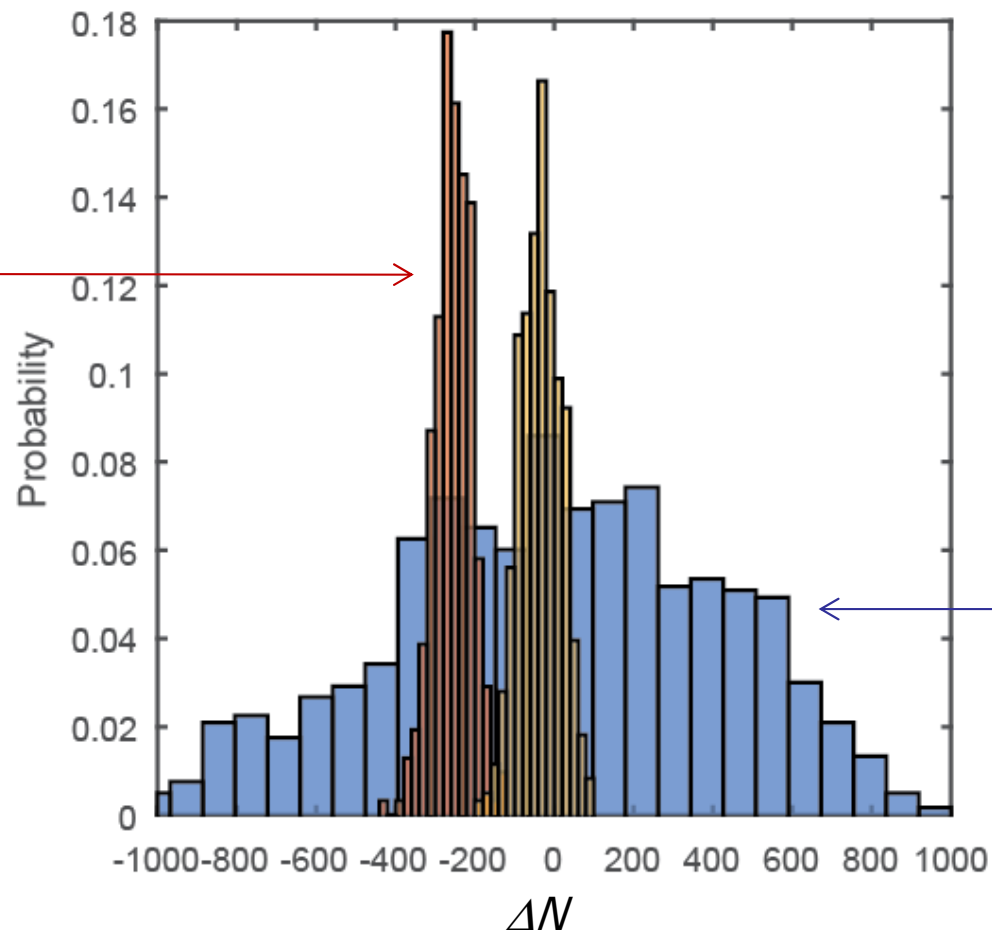
Consider  $N \sim 1e6$  atoms, each in a quantum superposition of two ground state energy levels.

Measure probability of finding atoms in one of these states:

*Entangled atoms*

*Reduced  
read-out noise*

*This data: 20 dB  
variance  
reduction*



*Uncorrelated  
atoms*

*"Shot-noise"  
Coin-toss  
statistics*

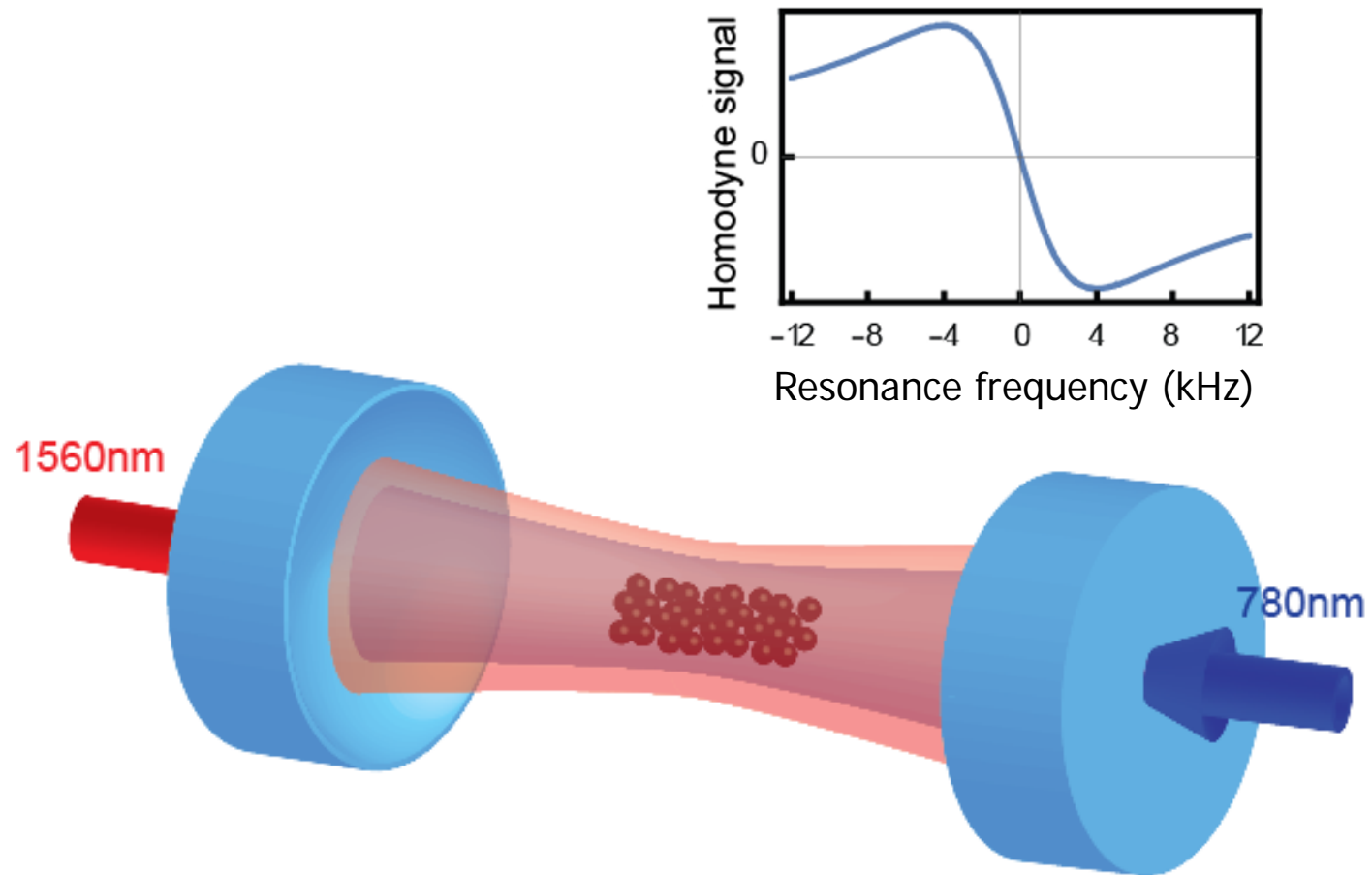
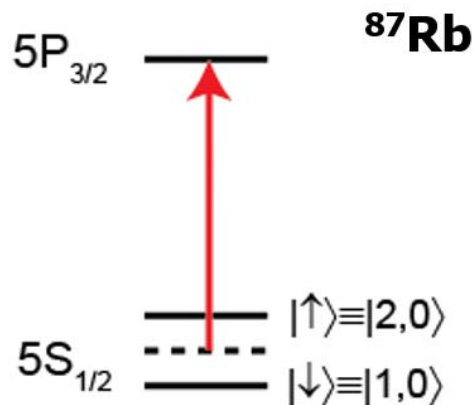




# Cavity-assisted entanglement

Dispersive atom-cavity interactions are used to realize a quantum non-demolition measurement of atom number.

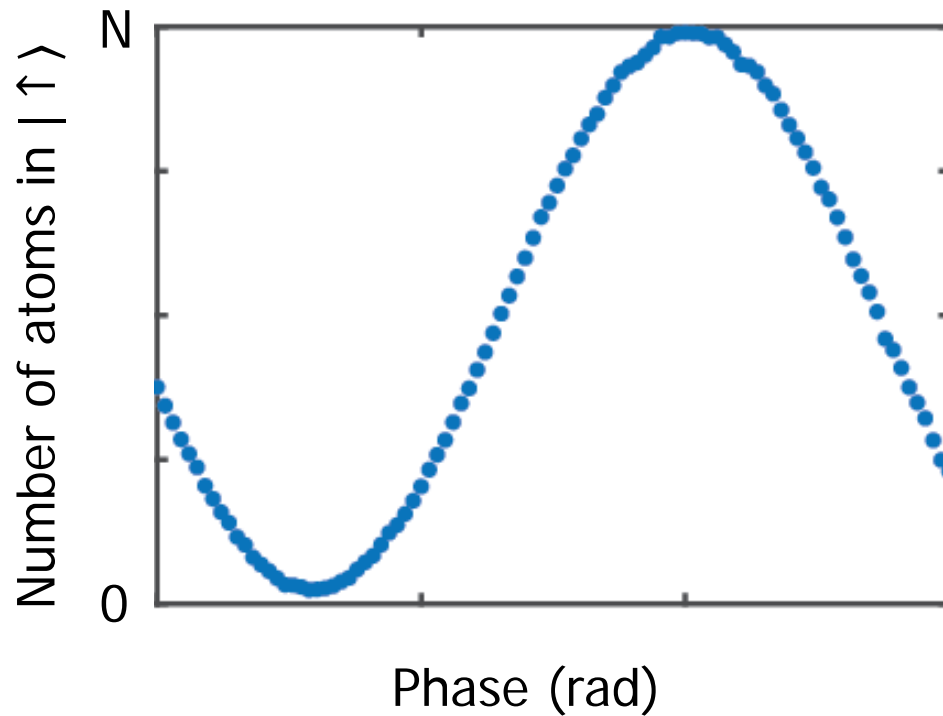
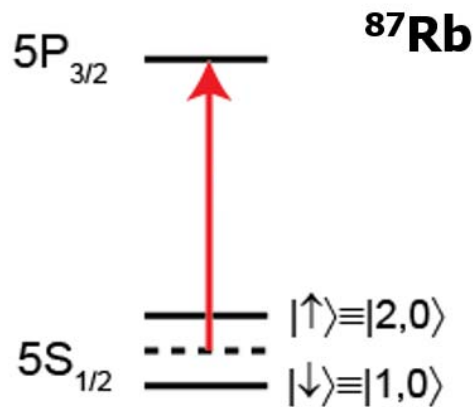
Measurement results in a metrologically useful many-atom entangled state.



# Coherence

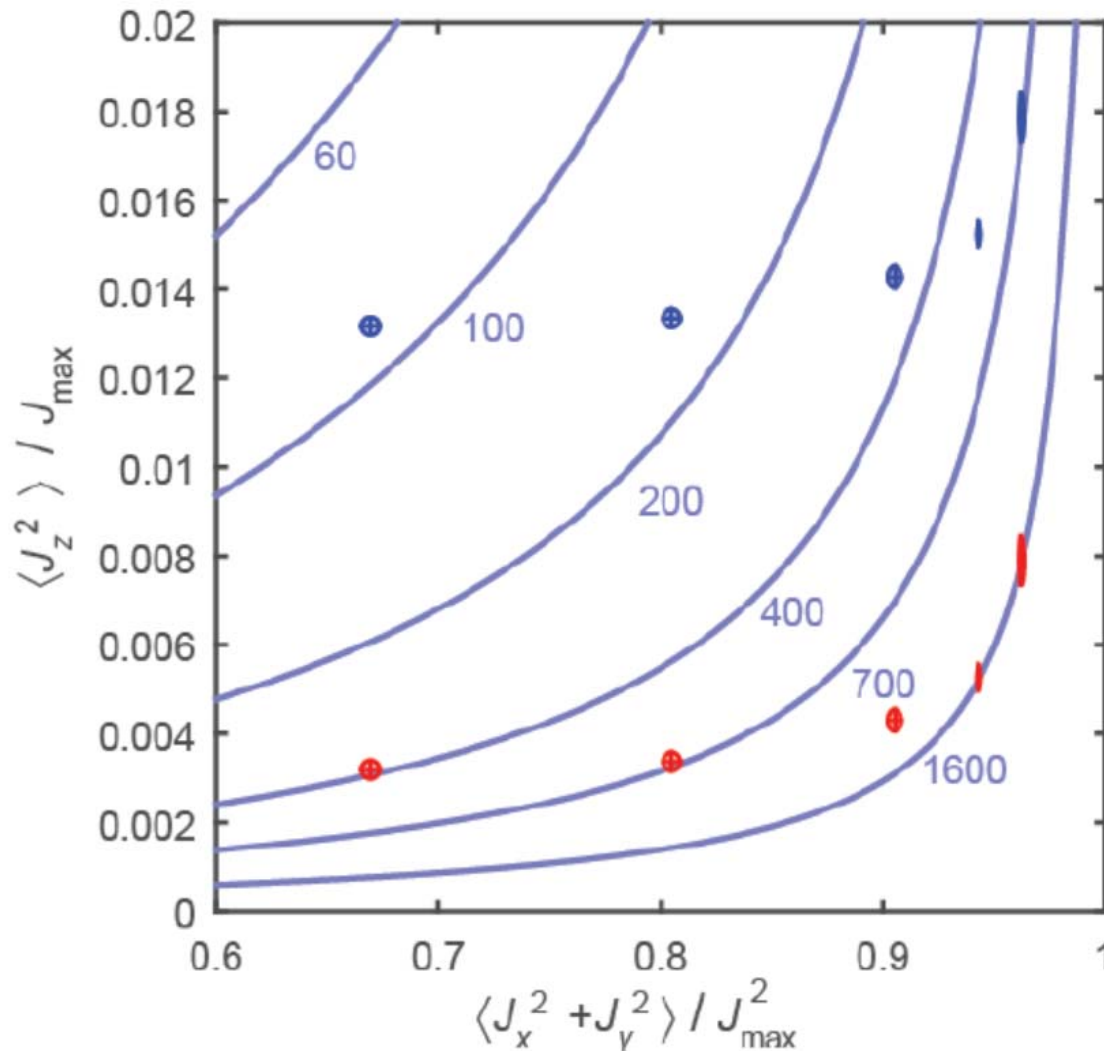
Measurement sequence:

- 1) Prepare atoms in  $|\downarrow\rangle$  state
- 2) Excitation (microwave) to  $(|\downarrow\rangle + |\uparrow\rangle)/\sqrt{2}$
- 3) Entangle/squeeze with optical (780 nm) probe;
- 4) Analysis pulse (microwave, scan excitation phase).



# Entanglement depth

States are entangled. By how much?



Red: theory assuming no coupling inhomogeneity, 1600 atoms entangled, following B. Lucke, PRL 112 (2014).

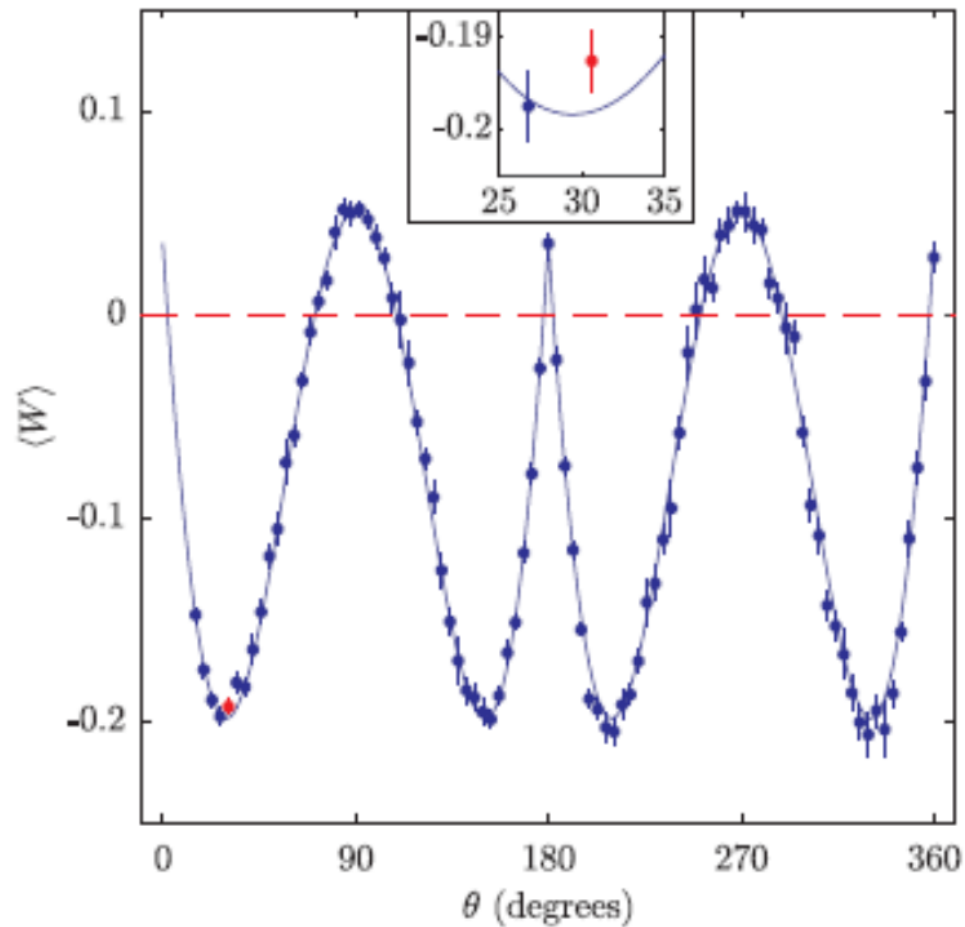
Blue: inferred coupling inhomogeneity, 680 atoms entangled.

# Bell witness

Bell witness for many-particle entangled states:

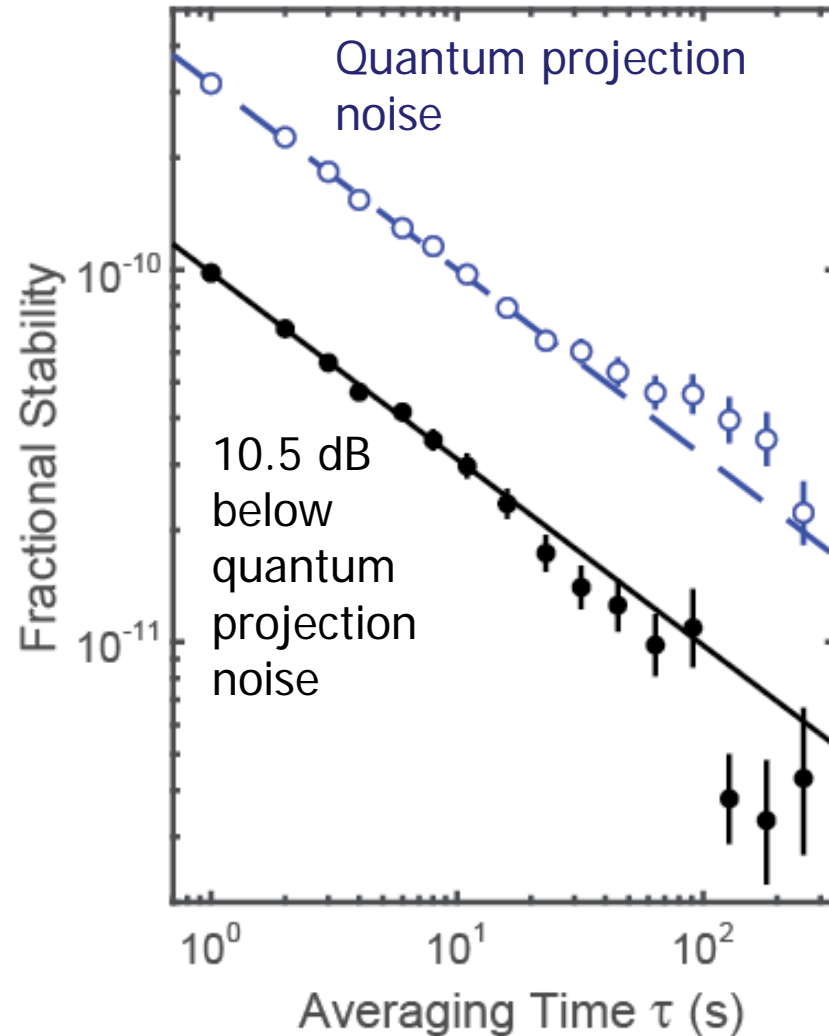
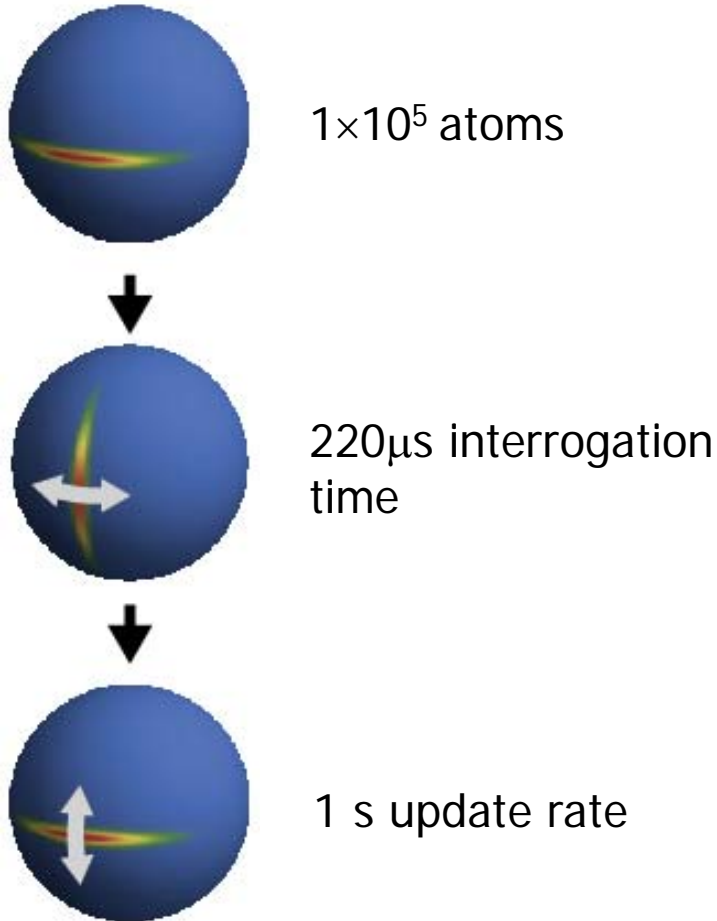
$$\langle W \rangle = -|\mathcal{J}_{1,n}| + (\mathbf{z} \cdot \mathbf{n})^2 \mathcal{J}_{2,z} + 1 - (\mathbf{z} \cdot \mathbf{n})^2 \geq 0$$

Engelsen, et al., PRL  
2017, following  
Schmied, et al., Science  
(2016)



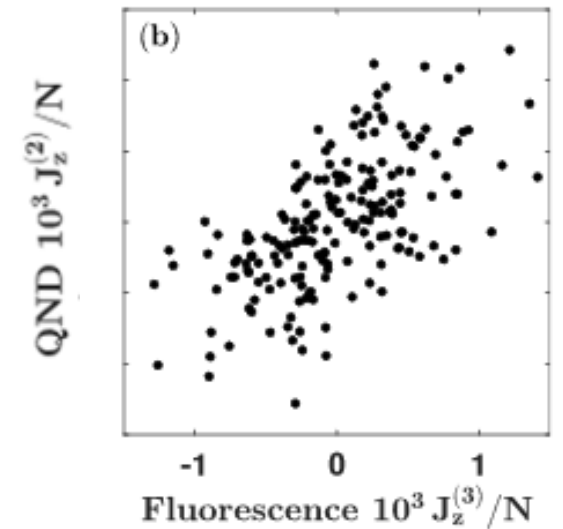
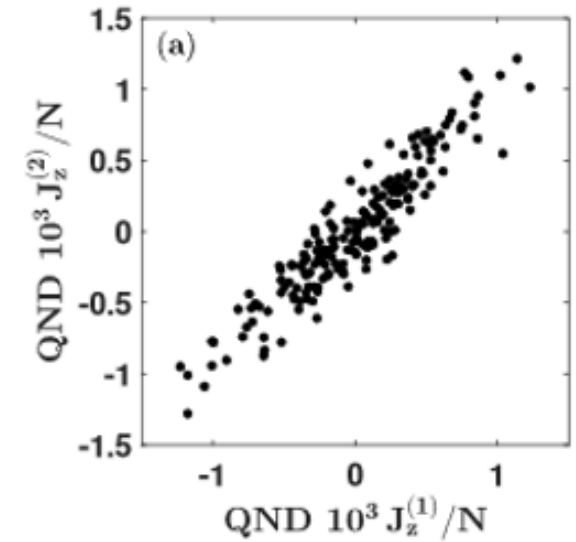
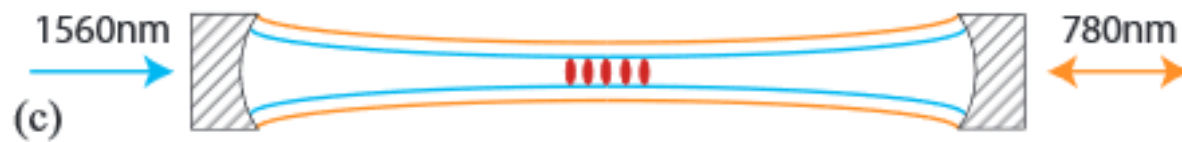
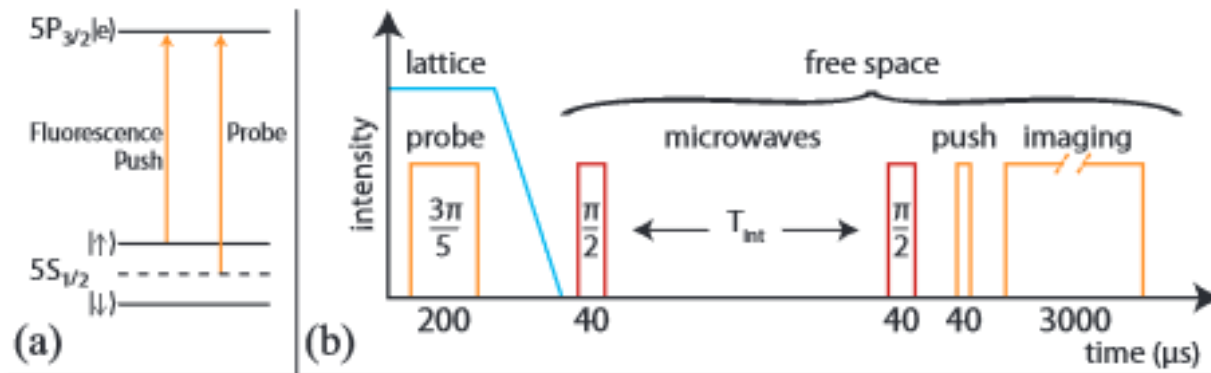
56  $\sigma$  violation of  
witness criterion

# Atomic clock implementation

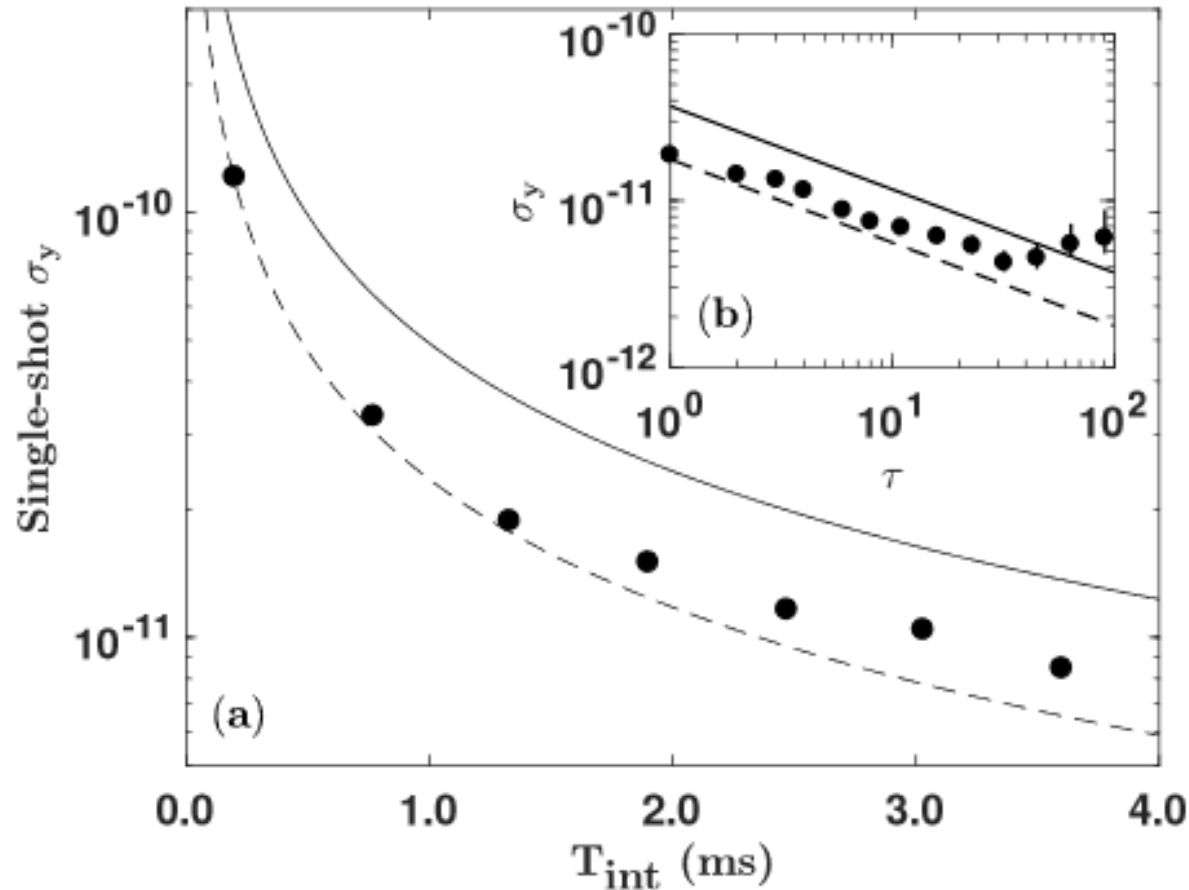


Limited by  $\mu$ -wave LO phase noise. Nature (2016)

# Fluorescence detection



# Entangled state, free-space, clock



*-6.4 dB below the classical (projection) noise limit*

*Limited by technical noise in read-out CCD*

Many atom entanglement persists over distances of mm and time scales of msec ....

# Comparison with Microsemi 5071A



*Spin-squeezed clocks  
are now comparable  
to Microsemi 5071A  
High Performance  
spec.*

## Stability (Allan Deviation)

Average Time (s)	Standard Performance	High Performance
0.01	$\leq 7.5 \times 10^{-11}$	$\leq 7.5 \times 10^{-11}$
0.1	$\leq 1.2 \times 10^{-11}$	$\leq 1.2 \times 10^{-11}$
1	$\leq 1.2 \times 10^{-11}$	$\leq 5.0 \times 10^{-12}$
10	$\leq 8.5 \times 10^{-12}$	$\leq 3.5 \times 10^{-12}$
100	$\leq 2.7 \times 10^{-12}$	$\leq 8.5 \times 10^{-13}$



# Future

Massively entangled states  
interfering over meter scales

Applications to precision  
gravitational physics and geodesy



# Thanks

Peter Asenbaum

Yunfan Wu

Ben Pichler

Julian Martinez

Chris Overstreet

Remy Notermans

Jason Hogan (Stanford)

Onur Hosten (IST, Vienna)

Tim Kovachy (Northwestern)

